

QFT in Curved Spacetime

Problems 3

1. By evaluating the integral

$$\int_{\mathbb{R}^n} d^n x e^{-x^2}$$

in two different ways, show that the area of the n-sphere, A_n , is given by the expression

$$A_n = \frac{2\pi^{(n+1)/2}}{\Gamma\left(\frac{n+1}{2}\right)} .$$

Let $G(x, x')$ denote the Feynman propagator for a massive scalar field. Thus

$$\langle 0|T(\varphi(x)\varphi(x'))|0\rangle = -iG(x, x') \quad \text{and} \quad (\square - m^2)G(x, x') = -g^{-1/2} \delta(x, x') .$$

By making use of the representation

$$\frac{1}{k^2 + m^2} = i \int_0^\infty ds e^{-is(k^2+m^2)} ,$$

(notice that this builds in the $i\epsilon$ prescription for m^2) show that, in Minkowski space, for the standard vacuum

$$G(x, x') = \frac{1}{2\pi} \left(\frac{m^2}{8\pi^2\sigma} \right)^{\frac{n-2}{4}} K_{\frac{n-2}{2}} \left(\sqrt{2m^2\sigma} \right)$$

where K_ν denotes the modified Bessel function and

$$\sigma = \frac{1}{2}(x - x')^2 + i\epsilon .$$

Take the limit $m^2 \rightarrow 0$ to show that, in the massless case,

$$G(x, x') = \frac{1}{8\pi^2\sigma} .$$

2. The Kontorovich-Lebedev transform (see Erdélyi et al. *Higher Transcendental Functions* vol II) is classically given by the pair of relations

$$f(\nu) = \int_0^\infty d\xi \tilde{f}(\xi) K_{i\nu}(\xi) ,$$

$$\tilde{f}(\xi) = \frac{2}{\pi^2 \xi} \int_0^\infty d\nu \nu \sinh(\pi\nu) K_{i\nu}(\xi) f(\nu) .$$

Show that these relations are equivalent to the relations

$$\int_0^\infty \frac{d\xi}{\xi} K_{i\nu}(\xi) K_{i\nu'}(\xi) = \frac{\pi^2}{2} \frac{\delta(\nu - \nu')}{\nu \sinh \nu\pi}$$

$$\int_0^\infty d\nu \nu \sinh \nu\pi K_{i\nu}(\xi) K_{i\nu}(\xi') = \frac{\pi^2}{2} \xi \delta(\xi - \xi') .$$

Show that the modes

$$v_{\mathbf{k}\nu}(x) = \frac{1}{2\pi^2} \sqrt{\sinh \nu\pi} e^{-i\nu\tau} K_{i\nu}(\mu\xi) e^{i\mathbf{k}\cdot\mathbf{y}}$$

are orthonormal with respect to the inner product

$$(v_{\mathbf{k}\nu}, v_{\mathbf{k}'\nu'}) = i \int d^2\mathbf{y} \int_0^\infty \frac{d\xi}{\xi} v_{\mathbf{k}\nu}^* \overset{\leftrightarrow}{\partial}_\tau v_{\mathbf{k}'\nu'} .$$

Expand a scalar field in modes and define a Feynman propagator appropriate to the modes we have just defined. Show that this Feynman propagator is given by the expression

$$G(x, x') = \int_0^\infty \frac{d\nu}{\pi^2} \sinh \nu\pi \int \frac{d^2\mathbf{k}}{(2\pi)^2} K_{i\nu}(\mu\xi) K_{i\nu}(\mu'\xi') e^{-i|\tau-\tau'|+i\mathbf{k}\cdot(\mathbf{y}-\mathbf{y}')} .$$