

QFT in Curved Spacetime

Problem Set 1

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1) Maxwell's equations :

$$g^{ac} \nabla_c F_{ab} = 0, \quad \nabla_{[a} F_{bc]} = 0$$

under $\tilde{g}_{ab} = \Omega^2 g_{ab}$ (so $\tilde{g}^{ab} = \Omega^{-2} g^{ab}$)

If ∇_a is associated with g_{ab} and

$\tilde{\nabla}_a$ is associated with \tilde{g}_{ab} , we denote
for some covector w_a

$$\tilde{\nabla}_a w_b = \nabla_a w_b - C^c{}_{ab} w_c \text{ so that}$$

$$C^c{}_{ab} = \frac{1}{2} \tilde{g}^{cd} \{ \nabla_a \tilde{g}_{bd} + \nabla_b \tilde{g}_{ad} - \nabla_d \tilde{g}_{ab} \}$$

since $\nabla_a \tilde{g}_{bc} = \nabla_a (\Omega^2 g_{bc}) = 2 \Omega g_{bc} \nabla_a \Omega$

$$\begin{aligned} \text{so } C^c{}_{ab} &= \Omega^{-1} g^{cd} \{ g_{bd} \nabla_a \Omega + g_{ad} \nabla_b \Omega - g_{ab} \nabla_d \Omega \} \\ &= 2 S^c{}_{(a} \nabla_{b)} \Omega - g_{ab} g^{cd} \nabla_d \Omega \end{aligned}$$

Note that $C^c{}_{ab} = C^c{}_{ba}$, symmetric in lower indices

Assume we are in n dimensions and F_{ab}
has scaling weight S (i.e. $F_{ab} \rightarrow \tilde{F}_{ab} = \Omega^S F_{ab}$)

$$\therefore g^{ac} \nabla_c F_{ab} \rightarrow \tilde{g}^{ac} \tilde{\nabla}_c (\Omega^S F_{ab})$$

$$= \Omega^{-2} g^{ac} \{ \nabla_c (\Omega^S F_{ab}) - C^d{}_{ca} \Omega^S F_{db} - C^d{}_{cb} \Omega^S F_{ad} \}$$

$$= \Omega^{S-2} g^{ac} \nabla_c F_{ab} + (n-4+S) \Omega^{S-3} g^{ac} F_{ab} \nabla_c \Omega$$

$$\nabla_{[a} F_{bc]} \rightarrow \tilde{\nabla}_{[a} \Omega^S F_{bc]} \quad$$

$$\begin{aligned}
 &= \nabla_{[a} \Omega^S F_{bc]} - C_{(ab)}^d \underbrace{\nabla^S F_{d]c}}_{=0} - C_{(ac)}^d \underbrace{\nabla^S F_{b]d}}_{=0} \\
 &= \Omega^S \nabla_{[a} F_{bc]} + S \Omega^{S-1} (\nabla_{ca} \Omega) F_{bc]} \quad \text{symmetrise then} \\
 &\qquad\qquad\qquad \text{antisymmetrise}
 \end{aligned}$$

$$\text{so in } \underline{n=4} \quad \underline{S=0}$$

$$\begin{aligned}
 g^{ac} \nabla_c F_{ab} &\rightarrow \Omega^2 g^{ac} \nabla_c F_{ab} \\
 \nabla_{[a} F_{bc]} &\rightarrow \nabla_{[a} F_{bc]}
 \end{aligned}
 \quad \left. \right\} \Rightarrow \begin{array}{l} \text{conformally} \\ \text{invariant} \end{array} \quad R$$

* 1) on the level of action, 4-dim

$$S = \int d^4x \sqrt{-g} \left(-\frac{1}{4} F_{ab} F^{ab} \right)$$

$$= \int d^4x \sqrt{-g} \left(-\frac{1}{4} \right) g^{ac} g^{bd} F_{ab} F_{cd}$$

for $g_{ab} \rightarrow \Omega^2 g_{ab}$ ∵ inverse $g^{ab} \rightarrow \Omega^{-2} g^{ab}$

$$g = \det(g_{ab}) \rightarrow \det(\Omega^2 g_{ab}) = (\Omega^2)^4 \det(g_{ab})$$

$$= \Omega^8 g$$

$$\therefore \sqrt{-g} \rightarrow \Omega^4 \sqrt{-g}$$

$$\therefore S \rightarrow -\frac{1}{4} \int d^4x \Omega^4 \sqrt{-g} \Omega^{-2} g^{ac} \Omega^{-2} g^{bd} F_{ab} F_{cd}$$

$$= \int d^4x \sqrt{-g} \left(-\frac{1}{4} F_{ab} F^{ab} \right) = S$$

∴ Action of Maxwell theory is invariant.

2) stress-energy tensor is defined by

$$T^{ab} = -\frac{2}{\sqrt{g}} \frac{\delta S}{\delta g_{ab}} \Rightarrow S = -\frac{1}{2} \int d^4x \sqrt{g} T^{ab} \delta g_{ab}$$

infinitesimally $g_{ab} \rightarrow \eta^{ab} g_{ab}$ becomes

$$g_{ab} \rightarrow (1+\epsilon) g_{ab} \quad \text{for small parameter } \epsilon(x)$$

$$\therefore \delta g_{ab} = \epsilon g_{ab} \Rightarrow S = -\frac{1}{2} \int d^4x \sqrt{g} T^a_a \epsilon(x)$$

i. under conformal transformation $\delta S = 0$

$$\therefore \int d^4x \sqrt{g} T^a_a \epsilon(x) = 0$$

true for arbitrary small $\epsilon(x)$

$$\Rightarrow \underbrace{T^a_a}_{} = 0 \quad \text{true free.}$$

* 2) stress energy tensor $T^{ab} = T^{ba}$

conservation $\nabla_a T^{ab} = 0$ under conformal transformation ($T^{ab} \rightarrow \tilde{T}^{ab} = \Omega^s T^{ab}$, $\tilde{g}_{ab} = \Omega^2 g_{ab}$) in n dimensions:

$$\begin{aligned}\rightarrow \quad & \tilde{\nabla}_a (\Omega^s T^{ab}) \\ = & \nabla_a (\Omega^s T^{ab}) + C^a_{ac} \Omega^s T^{cb} + C^b_{ac} \Omega^s T^{ac} \\ = & \Omega^s \nabla_a T^{ab} + cst + (2s+1) \Omega^{s-1} T^{ab} \nabla_a \Omega - \Omega^{s-1} g^{ba} T \nabla_a \Omega\end{aligned}$$

($T = T^a_a$ = trace of T)

choose scale weight $s = -n-2$

For $n=4$, $s=-6$

$$\nabla_a T^{ab} \rightarrow \Omega^{-6} \nabla_a T^{ab} - \Omega^{-7} g^{ba} (\nabla_a \Omega) T$$

conformally invariant only if $T=0$

$\Leftrightarrow T^{ab}$ trace free

3) for $d=n$ dimensional wave equation.

$$(\text{massless}) \Rightarrow (\square + \alpha R)\phi = 0$$

$$\text{under } \tilde{g}_{ab} = \Omega^2 g_{ab}, \phi \rightarrow \Omega^{\frac{s}{2}} \phi$$

$$\begin{aligned}\square \rightarrow \tilde{\square} &= \tilde{g}^{ab} \tilde{\square}_a \tilde{\square}_b \phi = \Omega^{-2} g^{ab} \tilde{\square}_a [\nabla_b (\Omega^{\frac{s}{2}} \phi)] \\ &= \Omega^{-2} g^{ab} [\nabla_a \nabla_b (\Omega^{\frac{s}{2}} \phi) - C_{ab} \nabla_c (\Omega^{\frac{s}{2}} \phi)] \\ &= \Omega^{s-2} g^{ab} \nabla_a \nabla_b \phi + (2s+n-2) \Omega^{s-3} g^{ab} \nabla_a \Omega \nabla_b \phi \\ &\quad + s \Omega^{s-3} \phi g^{ab} \nabla_a \nabla_b \Omega \\ &\quad + s(n+s-3) \Omega^{s-4} \phi g^{ab} \nabla_a \Omega \nabla_b \Omega\end{aligned}$$

$$\begin{aligned}R_{abc}{}^d \rightarrow \tilde{R}_{abc}{}^d &= R_{abc}{}^d - 2 \nabla_{[a} C^e{}_{b]c} + 2 C^e{}_{c[a} C^d{}_{b]e} \\ &= R_{abc}{}^d + 2 \delta^d{}_{[a} \nabla_{b]} \nabla_{c]n} \Omega - 2 g^{de} g_{c[e} \nabla_{b]} \nabla_{f]n} \Omega \\ &\quad + 2 (\nabla_{[a} \nabla_{b]} \Omega) \delta^d{}_{[b]} \nabla_{c]n} \Omega - 2 (\nabla_{[a} \nabla_{b]} \Omega) g_{b]c} g^{df} \nabla_f \nabla_{ln} \Omega \\ &\quad - 2 g_{c[a} \delta^d{}_{b]} g^{ef} (\nabla_{ln} \Omega) \nabla_f \nabla_{ln} \Omega\end{aligned}$$

$$\begin{aligned}\text{So } \tilde{R}_{ac} &= R_{ac} - (n-2) \nabla_a \nabla_{cn} \Omega - g_{ac} g^{de} \nabla_d \nabla_e \Omega \\ &\quad + (n-2) (\nabla_{[a} \nabla_{b]} \Omega) \nabla_{c]n} \Omega - (n-2) g_{ac} g^{de} (\nabla_d \nabla_e \Omega) \nabla_{ln} \Omega\end{aligned}$$

Contracting with \tilde{g}^{ac}

$$\Rightarrow \tilde{R} = \Omega^{-2} (R - 2(n-1)g^{ac}\nabla_a\nabla_c \ln \Omega - (n-2)(n-1)g^{ac}(\nabla_a \ln \Omega)\nabla_c \ln \Omega)$$

choose $a = -\frac{n-2}{4(n-1)}$, $S = 1 - \frac{n}{2}$, then

$$(\square - \frac{n-2}{4(n-1)})\phi \rightarrow (\tilde{g}^{ab}\tilde{\nabla}_a\tilde{\nabla}_b - \frac{n-2}{4(n-1)}\tilde{R})/\Omega^{1-\frac{n}{2}}\phi$$

$$= \Omega^{-1-\frac{n}{2}} \left(g^{ab} \nabla_a \nabla_b - \frac{n-2}{4(n-1)} R \right) \phi$$

\rightarrow conformally invariant

scale weight (for $n=d$), is

$$S = 1 - \frac{d}{2}$$

$$\overbrace{\hspace{10em}}^D$$

$$\text{For } d=n=4, \quad a = -\frac{4-2}{4(4-1)} = -\frac{1}{6}$$

$$\therefore \text{in } d=4, \quad (\square - \frac{R}{6})\phi \text{ is}$$

conformally invariant.

2

conventions: $\psi^A \epsilon_{AB} = \psi_B$, $\psi^A = \epsilon^{AB} \psi_B$

$$\text{identities: } \epsilon_{A'C'} \epsilon_{B'C'} = -\epsilon_{B'C'} \epsilon_{C'A'} = -\epsilon_{B'A'} \\ = \epsilon_{A'B'}$$

$$\epsilon_B^c \phi_{Ac} = -\epsilon_{Bc} \phi_A^c = \phi_A^c \epsilon_{cB} = \phi_{AB}$$

$$\epsilon_{A'C'} \bar{\phi}_{B'C'} = -\bar{\phi}_{B'C'} \epsilon_{C'A'} = -\bar{\phi}_{B'A'} = -\bar{\phi}_{A'B'}$$

$$\epsilon_{c'D'} \epsilon^{c'D'} = +2, \phi_{Ac} \phi_B^c = +\frac{1}{2} \epsilon_{AB} \phi_{cd} \phi^{cd} \text{ for symmetric } \bar{\phi}$$

$$\epsilon_{AC} \phi_B^c = -\phi_{AB}, \epsilon_{B'C'} \bar{\phi}_{A'C'} = \bar{\phi}_{A'B'}$$

These will be used in the calculation

$$\text{maxwell theory: } F_{ab} = \epsilon_{A'B'} \phi_{AB} + \epsilon_{AB} \bar{\phi}_{A'B'} \\ = F_{AA'BB'} \quad (\phi_{AB}, \bar{\phi}_{A'B'} \text{ symmetric})$$

$$T_{ab} = F_{ac} F_{b}^c - \frac{1}{4} g_{ab} F_{cd} F^{cd}$$

$$= -F_{AA'CC'} F_{BB'}^{CC'} + \frac{1}{4} \epsilon_{AB} \epsilon_{A'B'} F_{CC'DD'} F^{CC'DD'}$$

The quadratic terms:

$$F_{AA'CC'} F_{BB'}^{CC'} = (\epsilon_{A'C'} \phi_{Ac} + \epsilon_{AC} \bar{\phi}_{A'C'}) (\epsilon_{B'C'} \bar{\phi}_{B'C'} + \epsilon_{B'C'} \bar{\phi}_{B'C'}) \\ = \epsilon_{A'C'} \epsilon_{B'C'} \phi_{Ac} \phi_B^c + \epsilon_{A'C'} \epsilon_{B'C'} \phi_{Ac} \bar{\phi}_{B'C'} \\ + \epsilon_{AC} \epsilon_{B'C'} \bar{\phi}_{A'C'} \phi_B^c + \epsilon_{AC} \epsilon_{B'C'} \bar{\phi}_{A'C'} \bar{\phi}_{B'C'}$$

$$\begin{aligned}
&= \sum_{A'B'} \phi_{AC} \phi_B{}^C - \bar{\phi}_{A'B'} \phi_{AB} - \phi_{AB} \bar{\phi}_{A'B'} \\
&\quad + \sum_{AB} \bar{\phi}_{A'C'} \bar{\phi}_{B'C'} \\
&= -2\phi_{AB} \bar{\phi}_{A'B'} + \sum_{AB} \bar{\phi}_{A'C'} \bar{\phi}_{B'C'} + \sum_{A'B'} \bar{\phi}_{AC} \phi_B{}^C \\
&= -2\phi_{AB} \bar{\phi}_{A'B'} + \frac{1}{2} \sum_{AB} \sum_{A'B'} \bar{\phi}_{C'D'} \bar{\phi}^{C'D'} \\
&\quad + \frac{1}{2} \sum_{AB} \sum_{A'B'} \phi_{CD} \phi^{CD}
\end{aligned}$$

$$\begin{aligned}
&\frac{1}{4} F_{CC'DD'} F^{CC'DD'} \\
&= \frac{1}{4} (\sum_{c'b'} \phi_{cb} + \sum_{cb} \bar{\phi}_{c'b'}) (\sum^{c'b'} \phi^{cd} + \sum^{cd} \bar{\phi}^{c'b'}) \\
&= \frac{1}{4} (\underbrace{\sum_{c'b'} \sum^{c'b'} \phi_{cb} \phi^{cd}}_{+2} + \underbrace{\sum_{c'b'} \sum^{cd} \phi_{cd} \bar{\phi}^{c'b'}}_{=0, \text{ sym} \times \text{antisym}}) \\
&\quad + \underbrace{\sum_{cb} \sum^{cd} \bar{\phi}_{cb} \bar{\phi}^{c'b'}}_{=0, \text{ sym} \times \text{antisym}} \\
&= +\frac{1}{2} \phi_{cd} \phi^{cd} + \frac{1}{2} \bar{\phi}_{c'b'} \bar{\phi}^{c'b'} \\
\Rightarrow T_{ab} &= 2\phi_{AB} \bar{\phi}_{A'B'} + \frac{1}{2} \sum_{AB} \sum_{A'B'} (-\bar{\phi}_{C'D'} \bar{\phi}^{c'd'} \\
&\quad - \cancel{\phi_{ab} \phi^{cd}} + \cancel{\phi_{ab} \phi^{cd}} + \cancel{\bar{\phi}_{c'b'} \bar{\phi}^{c'b'}}) = \boxed{2\phi_{AB} \bar{\phi}_{A'B'}}
\end{aligned}$$

→ scalar massless wave equation with $a=0$

$$T_{ab} = \partial_a \phi \partial_b \phi - \frac{1}{2} g_{ab} \partial_c \phi \partial^c \phi$$

$$= D_{AA'} \phi D_{BB'} \phi - \frac{1}{2} \epsilon_{AB} \epsilon_{A'B'} D_{CC'} \phi D^{CC'} \phi$$

$$= D_{AA'} \phi D_{BB'} \phi - \epsilon_{A'B'} D_{AC'} \phi D_{B'C'} \phi$$

$$(\text{we used } D_{AC'} \phi D_{B'C'} \phi = +\frac{1}{2} \epsilon_{AB} D_{CC'} \phi D^{CC'} \phi)$$

and notice that $\epsilon_{A'B'} D_{AC'} \phi \epsilon^{C'D'} D_{BD'} \phi$

$$\text{for } \begin{matrix} A' = 1 & B' = 2 \\ 1 & 2 & A & 1 & 1 & 2 & B & 2 \\ 1 & 2 & A & 2 & 2 & 1 & B & 1 \end{matrix}$$

$$(\epsilon_{12} = +1 \quad \epsilon^{12} = +1 \quad \epsilon^{21} = -1)$$

$$\text{so this} = +D_{AA'} \phi D_{BB'} \phi - D_{AB'} \phi D_{BA'} \phi$$

$$\therefore T_{ab} = D_{AA'} \phi \cancel{D_{BB'} \phi} - (+D_{AA'} \phi D_{BB'}) \\ - D_{AB'} \phi D_{BA'} \phi$$

$$\Rightarrow T_{ab} = \underbrace{D_{AB'} \phi D_{BA'} \phi}_{}$$

- when contracted with $\ell^a \ell^b$ with ℓ being time-like, we can write $\phi_{AB} = \alpha_A \alpha_B$ for some α

$$\text{and } \ell^a = f^{AA'} = \sigma^A \bar{\sigma}^{A'} \text{ for some } \sigma$$

$$\text{so with } T_{ab} = \phi_{AB} \bar{\phi}_{A'B'}$$

$$\begin{aligned} T_{ab} \ell^a \ell^b &= \phi_{AB} \bar{\phi}_{A'B'} \sigma^A \bar{\sigma}^{A'} \sigma^B \bar{\sigma}^{B'} \\ &= \alpha_A \alpha_B \bar{\alpha}_{A'} \bar{\alpha}_{B'} \sigma^A \bar{\sigma}^{A'} \sigma^B \bar{\sigma}^{B'} \\ &= |\alpha_A \sigma^A \alpha_B \sigma^B|^2 \geq 0 \end{aligned}$$

$$\text{Similarly, with } T^{ab} = \nabla_{AB'} \phi \nabla_{BA'} \phi$$

and ϕ is real scalar

$$\begin{aligned} T_{ab} \ell^a \ell^b &= \nabla_{AB'} \phi \nabla_{BA'} \phi \sigma^A \bar{\sigma}^{A'} \sigma^B \bar{\sigma}^{B'} \\ &= |\sigma^A \nabla_{AB'} \phi \bar{\sigma}^{B'}|^2 \geq 0 \end{aligned}$$

Hence the dominant energy condition

D.

[4] i) $\nabla_l = \ell^a \nabla_a$ for null vector ℓ^a .

we know $\nabla_\ell \xi = -p\xi - \sigma\bar{\xi}$, $x^a = \xi m^a + \bar{\xi} \bar{m}^a$

For null congruences x^a , $\nabla_\ell x^a = x^b \nabla_b \ell^a$

$$\text{so } \nabla_\ell \nabla_\ell x^d = \nabla_\ell (x^b \nabla_b \ell^d)$$

$$= \nabla_\ell \nabla_x \ell^d = (\nabla_x \nabla_\ell + \nabla_{[\ell, x]}) \ell^d + R_{abc}{}^d \ell^a x^b \ell^c$$

\hookrightarrow Lie bracket = 0

$$= x^a \nabla_a (\nabla_\ell \ell^d) + \underbrace{R_{abc}{}^d \ell^a x^b \ell^c}_{=0 \because \text{geodesic equation}}$$

$$\Rightarrow \nabla_\ell \nabla_\ell x^d = R_{abc}{}^d \ell^a x^b \ell^c \text{ (geodetic deviation)}$$

Given that $\Phi_0 = C_{abcd} \ell^a m^b \ell^c m^d$
 \hookrightarrow Weyl curvature tensor

$$\text{and } \Phi_{00} = -\frac{1}{2} R_{ab} \ell^a \ell^b$$

$$x^a = \xi m^a + \bar{\xi} \bar{m}^a, m^a m_a = 0, \bar{m}^a \bar{m}_a = 0,$$

$$\nabla_\ell m^a = 0, \nabla_\ell \bar{m}^a = 0,$$

$$m^a \bar{m}_a = \partial^A \bar{l}^{A'} \bar{\partial}_{A'} l_A = \underbrace{|\partial^A l_A|}_1^2 = 1$$

use these we have

$$\bar{m}^d (\nabla_1 \nabla_1 \xi) + m^d (\nabla_1 \nabla_1 \xi) \\ = R_{abc}{}^d \ell^a (\bar{m}^b g + m^b \bar{g}) \ell^c$$

contract with m_d gives

$$(\underbrace{m_d \bar{m}^d}_{\textcircled{1}}) (\nabla_1 \nabla_1 \xi) + \underbrace{m_d m^d}_{\textcircled{2}} (\nabla_1 \nabla_1 \xi) \\ = R_{abcd} \ell^a \ell^c \bar{m}^b m^d \xi + R_{abcd} \ell^a m^b \ell^c m^d \xi$$

$$\textcircled{1} = \underbrace{R_{abcd} \ell^a \ell^c \bar{m}^b m^d}_{R_{abcd} = R_{cdab}} \xi = \frac{1}{2} R_{ac} \ell^a \ell^c \xi \\ \textcircled{2} \quad \text{use } \bar{m}^b m^d = \frac{1}{2} g^{bd} \\ = - \left(-\frac{1}{2} R_{ac} \ell^a \ell^c \right) \xi = - \Phi_\infty \xi$$

$\textcircled{2}$ $\because \ell^a m^b \ell^c m^d$ is trace free (l, m both null) \therefore it only picks up the trace free part of R_{abcd} , which is $-C_{abcd}$

$$\therefore \textcircled{2} = -C_{abcd} \ell^a m^b \ell^c m^d \xi = -\Psi_\infty \xi$$

Hence $\nabla_1 \nabla_1 \xi = -\Phi_\infty \xi - \Psi_\infty \xi \quad (*)$

in matrix form, $\nabla_L \vec{g} = -\rho \vec{g} - \sigma \vec{g}$ and its conjugate becomes

$$\nabla_L \begin{pmatrix} \vec{g} \\ \vec{\bar{g}} \end{pmatrix} = - \underbrace{\begin{pmatrix} \rho & \sigma \\ \bar{\sigma} & \bar{\rho} \end{pmatrix}}_P \begin{pmatrix} \vec{g} \\ \vec{\bar{g}} \end{pmatrix}$$

$$\Rightarrow \nabla_L \vec{\bar{z}} = -P \vec{\bar{z}}$$

and (X) becomes (including its conjugate)

$$\nabla_L \nabla_L \vec{\bar{z}} = -Q \vec{\bar{z}} \quad \text{where}$$

$$Q = \begin{pmatrix} \mathbb{E}_{00} & \Psi_0 \\ \bar{\Psi}_0 & \bar{\mathbb{E}}_{00} \end{pmatrix}$$

$$\text{So } \nabla_L \nabla_L \vec{\bar{z}} = -Q \vec{\bar{z}} = \nabla_L (-P \vec{\bar{z}})$$

$$= -(\nabla_L P) \vec{\bar{z}} - \underbrace{P \nabla_L \vec{\bar{z}}}_{-P \vec{\bar{z}}} = \underbrace{(-\nabla_L P + P^2)}_{-P \vec{\bar{z}}} \vec{\bar{z}}$$

$$\Rightarrow \nabla_L P = P^2 + Q$$

$$\Rightarrow \begin{cases} \nabla_L \rho = \rho \bar{\rho} + \sigma \bar{\sigma} + \mathbb{E}_{00} \\ \nabla_L \sigma = (\rho + \bar{\rho}) \sigma + \sigma \bar{\sigma} + \Psi_0 \end{cases} \quad (\text{Sach's equation})$$

$$2) \text{ vanishing twist} \Rightarrow \rho = \bar{\rho} \Rightarrow \bar{\rho}\bar{\rho} = \rho^2$$

vanishing shear $\Rightarrow \sigma = 0$, and that
 $\ell^\alpha \partial_\alpha = \partial_s = \frac{d}{ds}$ for some
affine s .

$$\text{flat space} \Rightarrow \underline{\underline{\mathcal{L}_\infty = 0}}, \underline{\underline{\mathcal{D}_\ell = \ell^\alpha \partial_\alpha}}$$

$$\text{so } \underline{\underline{\mathcal{D}_\ell \rho = \rho^2}} \Rightarrow \underline{\underline{\frac{d\rho}{ds} = \rho^2}}$$

$$\therefore \int_{\rho=\rho_0}^{\rho} \frac{d\rho'}{\rho'^2} = \int_{s=0}^s ds' \Rightarrow \frac{1}{\rho_0} - \frac{1}{\rho} = s$$

$$\therefore \frac{1}{\rho} = \frac{1}{\rho_0} - s = \frac{1 - \rho_0 s}{\rho_0}$$

$$\Rightarrow \rho = \frac{\rho_0}{1 - \rho_0 s} \quad \square$$

$\rho = \text{Re}(\rho)$ is the contraction, so the area
of horizon shrinks faster and faster ($\rho > 0$)
and the rate of shrinking goes to ∞
at time $s = \frac{1}{\rho_0}$

3) Assume still vanishing twist and shear, but in curved space so

$$\mathcal{E}_{\infty} \neq 0, \quad \mathcal{E}_{\infty} = -\frac{1}{2} R_{ab} \ell^a \ell^b.$$

But Einstein equation gives

$$R_{ab} - \frac{1}{2} g_{ab} (R - 2\Lambda) = -8\pi G T_{ab}$$

$$\therefore R_{ab} \ell^a \ell^b = -\frac{1}{2} (R - 2\Lambda) \ell^a \ell^a = -8\pi G T_{ab} \ell^a \ell^b$$

for future pointing null ℓ^a , $\ell^a \ell_a = 0$

dominant energy condition $T_{ab} \ell^a \ell^b \geq 0$

$$\therefore R_{ab} \ell^a \ell^b = -8\pi G T_{ab} \ell^a \ell^b \leq 0$$

$$\Rightarrow -2\mathcal{E}_{\infty} \leq 0 \quad \therefore \mathcal{E}_{\infty} \geq 0$$

Sach's equation in this case is

$$\frac{dp}{ds} = p^2 + \mathcal{E}_{\infty} \quad ! \quad \mathcal{E}_{\infty} \geq 0 \quad \therefore \frac{dp}{ds} \geq p^2$$

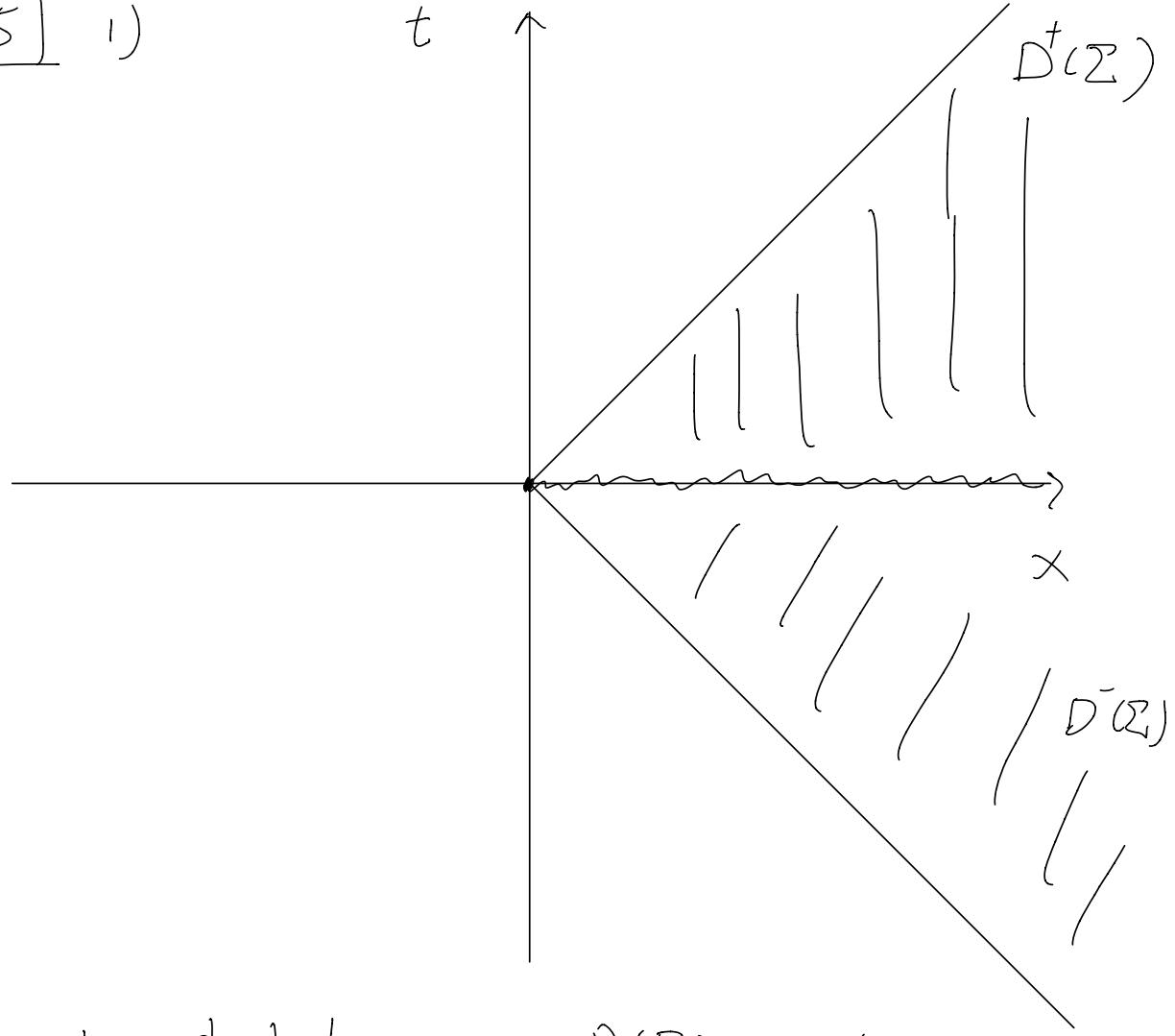
we see previously that if $\frac{dp}{ds} = p^2$ then with $p_0 > 0$, p blows up at proper time

$s = \frac{1}{p_0} = \text{finite}$, since with curvature

$\frac{dp}{ds} \geq p^2$, so p grows faster than
in the $\frac{dp}{ds} = p^2$ case, and blows up
within proper time $\frac{1}{p_0}$.

\Rightarrow blows up in finite time.

5 1)



The shaded area $D(\Sigma) = D^+(\Sigma) \cup D^-(\Sigma)$ is the Domain of dependence of hyper-surface $\Sigma: t=0, x>0$ in Minkowski spacetime.

(wavy line represents Σ)

$$2) \quad ds^2 = a^2 \gamma^2 dt^2 - d\gamma^2 - dy^2 - dz^2$$

consider coordinates change

$$\boxed{\begin{aligned} t &= \gamma \sinh(\alpha \tau) \\ x &= \gamma \cosh(\alpha \tau) \end{aligned}} \quad (*)$$

$$\Rightarrow dt = \sinh(\alpha \tau) d\gamma + a \gamma \cosh(\alpha \tau) d\tau$$

$$dx = \cosh(\alpha \tau) d\gamma + a \gamma \sinh(\alpha \tau) d\tau$$

$$\begin{aligned} \therefore dt^2 - dx^2 &= \sinh^2(\alpha \tau) d\gamma^2 + a^2 \gamma^2 \cosh^2(\alpha \tau) d\tau^2 \\ &\quad - \cosh^2(\alpha \tau) d\gamma^2 - a^2 \gamma^2 \sinh^2(\alpha \tau) d\tau^2 \\ &\quad + \cancel{2a \gamma \sinh(\alpha \tau) \cosh(\alpha \tau) d\gamma d\tau} \\ &\quad - \cancel{2a \gamma \sinh(\alpha \tau) \cosh(\alpha \tau) d\gamma d\tau} \end{aligned}$$

$$= (a^2 \gamma^2 d\tau^2 - d\gamma^2) (\underbrace{\cosh^2(\alpha \tau) - \sinh^2(\alpha \tau)}_1)$$

$$= a^2 \gamma^2 d\tau^2 - d\gamma^2$$

$$\text{so } ds^2 = dt^2 - dx^2 - dy^2 - dz^2 \text{ under } (*)$$

$$\therefore \Rightarrow \underbrace{\text{Minkowski}}$$

3) The domain of dependence above has boundary:
 $x = t\tau \Rightarrow x^2 = t^2$

$$\because \begin{cases} x = \zeta \cosh(a\tau) \\ t = \zeta \sinh(a\tau) \end{cases} \Rightarrow \zeta^2 \cosh^2(a\tau) = \zeta^2 \sinh^2(a\tau)$$

$$\Rightarrow \zeta^2(1 + \sinh^2(a\tau)) = \cancel{\zeta^2 \sinh^2(a\tau)}$$

$$\Rightarrow \zeta^2 = 0 \quad \therefore \quad \underline{\underline{\zeta = 0}}$$

The domain of dependence (interior) is

$x^2 > t^2, x > 0$, this corresponds

to $\underline{\zeta > 0}$

$\therefore D(\mathcal{Z})$ has been mapped to the half plane $\underline{\zeta \geq 0}$

4) for $(\xi, \gamma, z) = \text{const}$, $d\xi = dy = dz = 0$

$$\therefore d\xi^2 = a^2 \gamma^2 d\tau^2 \Rightarrow \frac{ds}{d\tau} = \pm a \gamma$$

there is gravitational redshift between different ξ coordinates, so a is the gravitational acceleration

5) Milne universe :

$$ds^2 = d\tau^2 - \tau^2 dx^2 - dy^2 - dz^2$$

consider $\boxed{\begin{array}{l} x = \tau \sinh(\chi) \\ t = \tau \cosh(\chi) \end{array}}$, we have

$$\begin{aligned} dt^2 - dx^2 &= (\tau \cdot \cosh(\chi) + \tau \sinh(\chi) \cdot d\chi)^2 \\ &\quad - (\tau \sinh(\chi) + \tau \cosh(\chi) \cdot d\chi)^2 \\ &= (\tau^2 - \tau^2 dx^2) (\underbrace{\cosh^2(\chi) - \sinh^2(\chi)}_{=1}) \\ &= d\tau^2 - \tau^2 dx^2 \end{aligned}$$

∴ the metric

$$ds^2 = dt^2 - dx^2 - dy^2 - dz^2 \Rightarrow \text{Minkowski (Flat)}$$

It is Globally hyperbolic