

To: Will Potter

55/55 100%

BS Problem Set 3

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Excellent Ziyan!

1. a) $c=1$

$$-dT^2 = -B(r,t)dt^2 + A(r,t)dr^2 + r^2d\theta^2 + r^2\sin^2\theta d\phi^2 \\ \equiv g_{\mu\nu}dx^\mu dx^\nu$$

($g_{\mu\nu}$ only has diagonal terms)

$$\Gamma_{\mu\nu}^\lambda = \frac{1}{2g_{\lambda\lambda}} \left(\frac{\partial g_{\lambda\mu}}{\partial x^\nu} + \frac{\partial g_{\lambda\nu}}{\partial x^\mu} - \frac{\partial g_{\mu\nu}}{\partial x^\lambda} \right)$$

$$g_{tt} = -B \quad g_{rr} = A \quad g_{\theta\theta} = r^2 \quad g_{\phi\phi} = r^2\sin^2\theta$$

$$\therefore \Gamma_{rt}^r = \frac{1}{2g_{rr}} \left(\frac{\partial g_{rr}}{\partial x^t} \right)$$

$$= \frac{1}{2g_{rr}} \frac{\partial g_{rr}}{\partial t} = \frac{\dot{A}}{2A} \checkmark$$

$$\Gamma_{rr}^t = \frac{1}{2g_{tt}} \left(-\frac{\partial g_{rr}}{\partial t} \right)$$

$$= -\frac{1}{2(-B)} \dot{A} = -\frac{\dot{A}}{2B} \checkmark$$

$$\Gamma_{tt}^t = \frac{1}{2g_{tt}} \left(\frac{\partial g_{tt}}{\partial t} + \frac{\partial g_{tt}}{\partial t} - \frac{\partial g_{tt}}{\partial t} \right)$$

$$= \frac{1}{2g_{tt}} \frac{\partial g_{tt}}{\partial t} = \frac{1}{2g_{tt}} \frac{\partial g_{tt}}{\partial t}$$

$$= \frac{1}{2(-B)} (-\dot{B}) = \frac{\dot{B}}{2B} \checkmark$$

b)

$$R_{\mu\nu} = \frac{1}{2} \frac{\partial^2 \ln |g|}{\partial x^\mu \partial x^\nu} - \frac{\partial \Gamma_{\mu\nu}^\lambda}{\partial x^\lambda}$$

$$\begin{aligned} \frac{1}{2} \frac{\partial^2 \ln |g|}{\partial x^\mu \partial x^\nu} &= \frac{1}{2} \frac{\partial^2}{\partial x^\mu \partial x^\nu} (\ln |g_{tt} g_{rr} g_{\theta\theta} g_{\phi\phi}|) \\ &= \frac{1}{2} \frac{\partial^2}{\partial x^\mu \partial x^\nu} (\ln |g_{tt}| + \ln |g_{rr}| + \ln |g_{\theta\theta}| + \ln |g_{\phi\phi}|) \end{aligned}$$

$$\begin{aligned} \frac{\partial \Gamma_{\mu\nu}^\lambda}{\partial x^\lambda} &= \frac{\partial}{\partial t} \Gamma_{\mu\nu}^t + \frac{\partial}{\partial r} \Gamma_{\mu\nu}^r + \frac{\partial}{\partial \theta} \Gamma_{\mu\nu}^\theta + \frac{\partial}{\partial \phi} \Gamma_{\mu\nu}^\phi \\ &= \frac{1}{2} \frac{\partial}{\partial t} \left[\frac{1}{g_{tt}} \left(\frac{\partial g_{t\mu}}{\partial x^\nu} + \frac{\partial g_{t\nu}}{\partial x^\mu} - \frac{\partial g_{\mu\nu}}{\partial t} \right) \right] \end{aligned}$$

$$+ \frac{1}{2} \frac{\partial}{\partial r} \left[\frac{1}{g_{rr}} \left(\frac{\partial g_{r\mu}}{\partial x^\nu} + \frac{\partial g_{r\nu}}{\partial x^\mu} - \frac{\partial g_{\mu\nu}}{\partial r} \right) \right]$$

$$+ \frac{1}{2} \frac{\partial}{\partial \theta} \left[\frac{1}{g_{\theta\theta}} \left(\frac{\partial g_{\theta\mu}}{\partial x^\nu} + \frac{\partial g_{\theta\nu}}{\partial x^\mu} - \frac{\partial g_{\mu\nu}}{\partial \theta} \right) \right]$$

$$+ \frac{1}{2} \frac{\partial}{\partial \phi} \left[\frac{1}{g_{\phi\phi}} \left(\frac{\partial g_{\phi\mu}}{\partial x^\nu} + \frac{\partial g_{\phi\nu}}{\partial x^\mu} - \frac{\partial g_{\mu\nu}}{\partial \phi} \right) \right]$$

$$\begin{aligned} &= \frac{1}{2} \left(\frac{\partial}{\partial t} \left(\frac{1}{g_{tt}} \frac{\partial g_{t\mu}}{\partial x^\nu} \right) + \frac{\partial}{\partial r} \left(\frac{1}{g_{rr}} \frac{\partial g_{r\mu}}{\partial x^\nu} \right) + \frac{\partial}{\partial \theta} \left(\frac{1}{g_{\theta\theta}} \frac{\partial g_{\theta\mu}}{\partial x^\nu} \right) \right. \\ &\quad \left. + \frac{\partial}{\partial \phi} \left(\frac{1}{g_{\phi\phi}} \frac{\partial g_{\phi\mu}}{\partial x^\nu} \right) \right) \end{aligned}$$

$$+ \frac{1}{2} \left(\frac{\partial}{\partial t} \left(\frac{1}{g_{tt}} \frac{\partial g_{t\nu}}{\partial x^\mu} \right) + \frac{\partial}{\partial r} \left(\frac{1}{g_{rr}} \frac{\partial g_{r\nu}}{\partial x^\mu} \right) + \frac{\partial}{\partial \theta} \left(\frac{1}{g_{\theta\theta}} \frac{\partial g_{\theta\nu}}{\partial x^\mu} \right) \right.$$

$$\left. + \frac{\partial}{\partial \phi} \left(\frac{1}{g_{\phi\phi}} \frac{\partial g_{\phi\nu}}{\partial x^\mu} \right) \right)$$

$$- \frac{1}{2} \left(\frac{\partial}{\partial t} \left(\frac{1}{g_{tt}} \frac{\partial g_{\mu\nu}}{\partial t} \right) + \frac{\partial}{\partial r} \left(\frac{1}{g_{rr}} \frac{\partial g_{\mu\nu}}{\partial r} \right) + \frac{\partial}{\partial \theta} \left(\frac{1}{g_{\theta\theta}} \frac{\partial g_{\mu\nu}}{\partial \theta} \right) \right.$$

$$\left. + \frac{\partial}{\partial \phi} \left(\frac{1}{g_{\phi\phi}} \frac{\partial g_{\mu\nu}}{\partial \phi} \right) \right)$$

$\mu = \nu$

If $\mu \neq \nu$ $\therefore g_{\mu\nu} \neq 0$ only if $\mu = \nu$

$$\therefore \frac{\partial \Gamma_{\mu\nu}^{\lambda}}{\partial x^{\lambda}} = \frac{1}{2} \frac{\partial}{\partial x^{\mu}} \left(\frac{1}{g_{\mu\mu}} \frac{\partial g_{\mu\mu}}{\partial x^{\nu}} \right) + \frac{1}{2} \frac{\partial}{\partial x^{\nu}} \left(\frac{1}{g_{\nu\nu}} \frac{\partial g_{\nu\nu}}{\partial x^{\mu}} \right)$$

$$= \frac{1}{2} \frac{\partial^2}{\partial x^{\mu} \partial x^{\nu}} \ln |g_{\mu\mu}| + \frac{1}{2} \frac{\partial^2}{\partial x^{\nu} \partial x^{\mu}} \ln |g_{\nu\nu}|$$

$$= \frac{1}{2} \frac{\partial^2 \ln |g_{\mu\mu} g_{\nu\nu}|}{\partial x^{\mu} \partial x^{\nu}}$$

If $\mu = \nu$ then $g_{\mu\nu} = g_{\mu\mu}$

~~$$\therefore \frac{\partial \Gamma_{\mu\nu}^{\lambda}}{\partial x^{\lambda}} = \frac{1}{2} \frac{\partial}{\partial x^{\mu}} \left(\frac{1}{g_{\mu\mu}} \frac{\partial g_{\mu\mu}}{\partial x^{\nu}} \right) + \frac{1}{2} \frac{\partial}{\partial x^{\nu}} \left(\frac{1}{g_{\nu\nu}} \frac{\partial g_{\nu\nu}}{\partial x^{\mu}} \right)$$~~
~~$$\frac{1}{2} \frac{\partial}{\partial x^{\mu}} \left(\frac{1}{g_{\mu\mu}} \frac{\partial g_{\mu\mu}}{\partial x^{\nu}} \right)$$~~

\therefore If $\mu = t$ $\nu = r$

$$\frac{\partial \Gamma_{\mu\nu}^{\lambda}}{\partial x^{\lambda}} = \frac{\partial \Gamma_{tr}^{\lambda}}{\partial x^{\lambda}} = \frac{1}{2} \frac{\partial^2}{\partial r \partial t} \ln |g_{tt} g_{rr}|$$

$$= \frac{1}{2} \frac{\partial^2}{\partial r \partial t} (\ln |g_{tt}| + \ln |g_{rr}|)$$

$$\frac{1}{2} \frac{\partial^2}{\partial x^{\mu} \partial x^{\nu}} |g| = \frac{1}{2} \frac{\partial^2}{\partial r \partial t} (\ln |g_{tt}| + \ln |g_{rr}| + \ln |g_{\theta\theta}| + \ln |g_{\phi\phi}|)$$

$\therefore g_{\theta\theta}, g_{\phi\phi}$ independent of t

$$\therefore \frac{1}{2} \frac{\partial^2}{\partial x^{\mu} \partial x^{\nu}} = \frac{1}{2} \frac{\partial^2}{\partial r \partial t} (\ln |g_{tt}| + \ln |g_{rr}|)$$

$$\rightarrow - \frac{\partial T_{tr}^{\lambda}}{\partial x^{\lambda}} + \frac{1}{2} \frac{\partial^2 \ln|g|}{\partial r \partial t} = 0 \quad 4/4$$

c) ∴ b) ∴ ~~$R_{tr} = T_{tr}$~~ ($|g| = AB r^4 \sin^2 \theta$)

$$\therefore R_{tr} = T_{tr}^{\eta} T_{\eta}^{\lambda} - \frac{T_{tr}^{\eta}}{2} \frac{\partial \ln|g|}{\partial x^{\eta}}$$

$$= T_{tt}^t T_{tt}^t + T_{tt}^r T_{rr}^t + T_{tr}^t T_{rt}^r + T_{tr}^r T_{rr}^r$$

$$- \frac{T_{tr}^t}{2} \frac{\partial \ln|g|}{\partial t} - \frac{T_{tr}^r}{2} \frac{\partial \ln|g|}{\partial r}$$

$$= \frac{\dot{B}}{2B} \frac{B'}{2B} + \frac{B'}{2A} \frac{\dot{A}}{2B} + \frac{B'}{2B} \frac{\dot{A}}{2A} + \frac{\dot{A} A'}{2A 2A}$$

$$- \frac{1}{2} \frac{B'}{2B} \frac{1}{AB r^4 \sin^2 \theta} \frac{\partial}{\partial t} (AB r^4)$$

$$- \frac{1}{2} \frac{\dot{A}}{2A} \frac{1}{AB r^4 \sin^2 \theta} \frac{\partial}{\partial r} (AB r^4)$$

$$= \frac{\dot{B}}{2B} \frac{B'}{2B} + \frac{B'}{2A} \frac{\dot{A}}{2B} + \frac{B'}{2B} \frac{\dot{A}}{2A} + \frac{\dot{A} A'}{2A 2A}$$

$$- \frac{1}{2} \frac{B'}{2B} \frac{1}{AB r^4} (\dot{A} B + \dot{B} A)$$

$$- \frac{1}{2} \frac{\dot{A}}{2A} \frac{1}{AB r^4} (r^4 A' B + r^4 B' A + 4r^3 AB)$$

$$= \frac{\dot{B}}{2B} \frac{B'}{2B} + \frac{B'}{2A} \frac{\dot{A}}{2B} + \frac{B'}{2B} \frac{\dot{A}}{2A} + \frac{\dot{A} A'}{2A 2A}$$

$$- \frac{B'}{2B} \frac{\dot{A}}{2A} - \frac{B'}{2B} \frac{\dot{B}}{2B} - \frac{\dot{A} A'}{2A 2A} - \frac{\dot{A} B'}{2A 2B} - \frac{\dot{A}}{2A} \cdot \frac{1}{2} \cdot \frac{4r^3}{r^4}$$

$$= - \frac{\dot{A}}{rA}$$

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(C) 2015

$$\frac{10/10}{10/10}$$

$$= \frac{10/10 + 10/10}{10/10 + 10/10} = \frac{20/10}{20/10} = 1$$

$$\frac{10/10}{10/10} = \frac{10/10}{10/10}$$

$$= \frac{\dot{A}}{A} + \frac{\dot{B}}{B} + \frac{\dot{C}}{C} + \frac{\dot{D}}{D}$$

$$\frac{10/10}{10/10} = \frac{10/10}{10/10}$$

$$= \frac{\dot{A}}{A} + \frac{\dot{B}}{B} + \frac{\dot{C}}{C} + \frac{\dot{D}}{D}$$

$$\frac{10/10}{10/10} = \frac{10/10}{10/10}$$

$$\frac{10/10}{10/10} = \frac{10/10}{10/10}$$

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$$\frac{10/10}{10/10} = \frac{10/10}{10/10}$$

2. a)

Gravitational field equation should take the form

$$G_{\mu\nu} = C T_{\mu\nu}$$

$T_{\mu\nu}$, the stress energy tensor, is the source of the gravitational field

$G_{\mu\nu}$ must be linear in the second derivative of $g_{\mu\nu}$ so that poisson equation is recovered

→ Now for a linear theory, $g_{\mu\nu} = \eta_{\mu\nu} + h_{\mu\nu}$ ($|h_{\mu\nu}| \ll 1$)

then the LHS of field equation must be a linear combination of second order derivatives of $h_{\mu\nu}$

For scalar second order derivatives, we need to multiply by $\eta_{\mu\nu}$, the only available tensor except $h_{\mu\nu}$, to turn the term into a 2nd rank tensor.

Also, $\because T_{\mu\nu}$ is symmetric ✓

\therefore We need 2nd order derivatives in $h_{\mu\nu}$ that are symmetric

→ only possible terms are

$$\square h_{\mu\nu}, \partial_\mu \partial_\nu h, (\partial_\alpha \partial_\mu h^\alpha_\nu + \partial_\alpha \partial_\nu h^\alpha_\mu), \eta_{\mu\nu} \square h, \eta_{\mu\nu} \partial_\alpha \partial^\alpha h^{\mu\nu}$$

b)

Conservation of energy:

covariant derivative of $T_{\mu\nu}$ vanish

$$T_{\mu\nu;\nu} = 0$$

Now in ~~weak~~ weak field, we replace covariant derivative with ordinary partial derivative

$$\Rightarrow \partial^\nu T_{\mu\nu} = 0 \quad \checkmark$$

$$\therefore \partial^\nu \left(\square h_{\mu\nu} + \alpha (\partial_\rho \partial_\nu h^\rho_\mu + \partial_\rho \partial_\mu h^\rho_\nu) + \beta (\partial_\mu \partial_\nu h) \right. \\ \left. + \eta_{\mu\nu} (\gamma \square h + \delta \partial_\rho \partial_\lambda h^{\rho\lambda}) \right) = 0$$

$$\therefore \partial_\rho \partial^\rho \partial^\nu h_{\mu\nu} + \alpha \partial_\rho \partial_\nu \partial^\mu h^\rho_\nu + \alpha \partial_\rho \partial_\nu \partial^\mu h^\rho_\mu \\ + \beta \partial_\mu \partial^\nu \partial_\nu h + \gamma \eta_{\mu\nu} \partial_\rho \partial^\rho h + \delta \eta_{\mu\nu} \partial_\rho \partial_\lambda h^{\rho\lambda} = 0$$

→ consider $\partial_\rho \partial^\rho \partial^\nu h_{\mu\nu} + \alpha \partial_\rho \partial_\nu \partial^\mu h^\rho_\nu$

$$= \partial_\rho \partial^\rho \partial^\nu h_{\mu\nu} + \alpha \partial_\rho \partial_\nu \partial^\mu h^\rho_\nu \xrightarrow{\text{switching } \rho, \nu}$$

$$= \partial_\rho \partial^\rho (\partial^\nu h_{\mu\nu} + \alpha \eta^{\mu\kappa} \partial_\rho h_{\kappa\nu})$$

$$= \partial_\rho \partial^\rho (\partial^\nu h_{\mu\nu} + \alpha \partial^\kappa h_{\kappa\nu})$$

$$= \partial_\rho \partial^\rho (\partial^\nu h_{\mu\nu} + \alpha \partial^\nu h_{\mu\nu}) = (1+\alpha) \partial_\rho \partial^\rho (\partial^\nu h_{\mu\nu})$$

$$\xrightarrow{\kappa \rightarrow \nu} = (1+\alpha) \square (\partial^\nu h_{\mu\nu})$$

→

consider

$$\begin{aligned} & \alpha \partial_\rho \partial_\nu \partial^\mu h^\rho_\mu + \delta \underbrace{\eta_{\mu\nu}}_{\partial_\nu} \partial^\mu \partial_\rho \partial_\lambda h^{\rho\lambda} \\ &= \alpha \partial_\rho \partial_\nu \partial^\mu \eta_{\lambda\mu} h^{\rho\lambda} + \cancel{\delta \partial_\nu \partial^\mu} \delta \partial_\nu \partial_\rho \partial_\lambda h^{\rho\lambda} \\ &= \alpha \partial_\rho \partial_\nu \partial_\lambda h^{\rho\lambda} + \delta \partial_\nu \partial_\rho \partial_\lambda h^{\rho\lambda} \\ &= (\alpha + \delta) \partial_\rho \partial_\nu \partial_\lambda h^{\rho\lambda} \end{aligned}$$

→

consider

$$\begin{aligned} & \beta \partial_\mu \partial^\mu \partial_\nu h + \cancel{\delta \partial^\mu} \partial^\mu \eta_{\mu\nu} \gamma \partial h \\ &= \beta \partial_\mu \partial^\mu \partial_\nu h + \gamma \partial_\nu \partial_\mu \partial^\mu h \\ &= (\beta + \gamma) \partial_\mu \partial^\mu \partial_\nu h \end{aligned}$$

$$\begin{aligned} \therefore (1 + \alpha) \square(\partial^\mu h_{\mu\nu}) + (\beta + \gamma) \partial_\mu \partial^\mu \partial_\nu h + (\alpha + \delta) \partial_\rho \partial_\nu \partial_\lambda h^{\rho\lambda} \\ = 0 \quad \text{for all metric } h_{\mu\nu} \end{aligned}$$

$$\therefore \begin{cases} 1 + \alpha = 0 \\ \beta + \gamma = 0 \\ \alpha + \delta = 0 \end{cases} \Rightarrow \begin{cases} \alpha = -1 \\ \delta = 1 \\ \gamma = -\beta \end{cases} \quad \checkmark \quad 5/5$$

c)

$$\square h_{\mu\nu} - 2\partial_\rho \partial_\mu h^\rho_\nu + \beta \partial_\rho \partial_\nu h - \eta_{\mu\nu} (\beta \square h - \partial_\rho \partial_\lambda h^{\rho\lambda})$$

$$= c T_{\mu\nu} \quad (1)$$

Now raise ν and take trace ~~(set $\nu = \mu$)~~ \circ

$$\square h = 2\partial_\rho \partial_\nu h^{\rho\nu} + \beta \square h$$

multiply (1) by ~~$\eta^{\mu\nu}$~~ $\eta^{\nu\mu} \Rightarrow$

$$\therefore \square h^\mu_\nu = \partial_\rho \partial^\mu h^\rho_\nu$$

$$c T^\mu_\nu = \square h^\mu_\nu - 2\partial_\rho \partial_\nu h^{\rho\mu} + \beta \partial_\nu \partial^\rho h^\rho_\mu - \eta^{\nu\mu} \eta_{\rho\sigma} (\beta \square h - \partial_\rho \partial_\lambda h^{\rho\lambda})$$

take trace (set $k = \nu = \mu$)

$$\Rightarrow \square h - 2\partial_\rho \partial_\lambda h^{\rho\lambda} + \beta \partial_\nu \partial^\nu h - \delta^\mu_\mu (\beta \square h - \partial_\rho \partial_\lambda h^{\rho\lambda}) = c T$$

$$\Rightarrow \square h - 2\partial_\rho \partial_\lambda h^{\rho\lambda} + \beta \square h - 4\beta \square h + 4\partial_\rho \partial_\lambda h^{\rho\lambda} = c T$$

$$\Rightarrow (1 + \beta - 4\beta) \square h + (-2 + 4) \partial_\rho \partial_\lambda h^{\rho\lambda} = c T$$

$$\dots \quad 2\partial_\rho \partial_\lambda h^{\rho\lambda} = (3\beta - 1) \square h + c T$$

$$\therefore \partial_\rho \partial_\lambda h^{\rho\lambda} = \frac{3\beta - 1}{2} \square h + \frac{c}{2} T \quad \checkmark$$

$$\text{Now } T = T^\mu_\mu = T_0^0 + T_1^1 + T_2^2 + T_3^3$$

$$= \eta^{00} T_{00} + \eta^{11} T_{11} + \eta^{22} T_{22} + \eta^{33} T_{33}$$

$$= -T_{00} + T_{11} + T_{22} + T_{33} \quad \checkmark$$

$$\therefore T_{\mu\nu} = -P g_{\mu\nu} + \left(P + \frac{P}{c^2}\right) U_\mu U_\nu$$

$$\therefore U_0 = \frac{d(ct)}{dt}, \quad U_i = \frac{dx_i}{dt}$$

Newtonian limit, $d(ct) \gg dx_i \Rightarrow U_0 \gg U_i$

Also $P/c^2 \ll P$, $g_{\mu\nu} \approx \eta_{\mu\nu}$

~~$$\therefore T_{\mu\nu} = -P \eta_{\mu\nu} + P U_\mu U_\nu \approx P U_\mu U_\nu$$~~

~~$$T_{00} = P + P U_0 U_0$$~~

~~$$T_{ii} = P$$~~

$$\therefore \frac{P}{c^2} \ll P \quad \therefore T_{\mu\nu} \approx P U_\mu U_\nu \quad (\text{dust})$$

$$\therefore U_0 \gg U_i \quad \therefore T_{00} \gg T_{ii}$$

$$\therefore T \approx -T_{00} \quad \checkmark$$

$$\therefore \frac{\partial}{\partial x^h} p^\lambda = \frac{3\beta - 1}{2} \square h - \frac{c T_{00}}{2} \quad \checkmark$$

good

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$$T^{\mu\nu} = \rho U^\mu U^\nu \quad T^{\mu\nu} = (P + P) U^\mu U^\nu + P g^{\mu\nu}$$

$$= \begin{pmatrix} P/c^2 & 0 & 0 & 0 \\ 0 & P & 0 & 0 \\ 0 & 0 & P & 0 \\ 0 & 0 & 0 & P \end{pmatrix}$$

$$T^{\mu\nu} = \sum P_i U^\mu U^\nu$$

$U^0 U^0$ always positive \Rightarrow sum up
to have $T_{ii} \approx T_{00}$
need $U^i \approx U^0 \Rightarrow$ relativistic gas.
gives

d)

$$\square h_{\mu\nu} - (\partial_\rho \partial_\mu h^\rho_\nu + \partial_\rho \partial_\nu h^\rho_\mu) + \beta \partial_\mu \partial_\nu h - \eta_{\mu\nu} (\beta \square h - \partial_\rho \partial_\lambda h^{\rho\lambda}) = c T_{\mu\nu}$$

Set $\nu=0$, $\nu=0$

Newtonian static limit: $\frac{\partial}{\partial t} = 0$, $\gamma_{00} = -1$

$$\therefore \square h_{00} + \beta \square h - \partial_\rho \partial_\lambda h^{\rho\lambda} = c T_{00}$$

$$\therefore \partial_\rho \partial_\lambda h^{\rho\lambda} = \frac{3\beta - 1}{2} \square h - \frac{c T_{00}}{2} \quad (\text{Newtonian limit})$$

$$\therefore \square h_{00} + \beta \square h - \left(\frac{3\beta}{2} + \frac{1}{2}\right) \square h + \frac{c T_{00}}{2} = c T_{00}$$

$$\therefore \square h_{00} + \frac{1-\beta}{2} \square h = \frac{c}{2} T_{00}$$

$$\therefore \frac{\partial}{\partial t} = 0 \quad \therefore \square \rightarrow \nabla^2$$

$$\therefore \nabla^2 h_{00} + \frac{1-\beta}{2} \nabla^2 h = \frac{c}{2} T_{00} \quad (2)$$

In Newtonian limit, from Newtonian geodesic

$$\text{we have } h_{00} = -\frac{2\Phi}{c^2}, \quad \nabla^2 \Phi = 4\pi G \rho, \quad T_{00} = \rho c^2$$

$$\therefore \nabla^2 \Phi = 4\pi G \rho \quad \text{becomes}$$

$$-\frac{c^2}{2} \nabla^2 h_{00} = \frac{4\pi G}{c^2} T_{00} \Rightarrow \nabla^2 h_{00} = -\frac{8\pi G}{c^4} T_{00} \quad (3)$$

$$\therefore c=1 \quad \therefore \nabla^2 h_{00} = -8\pi G T_{00} \quad (3)$$

Compare ② with ③, we have that

$$\frac{1-\beta}{2} = 0, \quad \frac{C}{2} = \cancel{-8\pi G} - 8\pi G$$

$$\Rightarrow \underline{\underline{\beta = 1}}, \quad \underline{\underline{C = -16\pi G}} \quad \checkmark \quad 7/7$$

e) ~~The general~~

In notes:

(339) \rightarrow the linear wave

(340) \rightarrow the source term consists of $T_{\mu\nu}$ and $T_{\mu\nu}$, the stress energy contribution of the gravitational radiation itself

(348) \rightarrow the gauge transformation that keeps the ~~equation~~ wave equation invariant.

e)

The general field equation needs to be sourced by $T_{\mu\nu}$, and its LHS should be linear in 2nd order derivatives of $g_{\mu\nu}$ to recover in weak field to recover Newton's Laws.

\therefore $R_{\mu\nu}$ and $g_{\mu\nu} R$ are only 2 available tensors

$$\therefore \text{ need } aR_{\mu\nu} + b g_{\mu\nu} R = C' T_{\mu\nu}$$

$$\text{raise both } \mu, \nu \Rightarrow aR^{\mu\nu} + b g^{\mu\nu} R = C T^{\mu\nu}$$

$$\text{Covariant derivative, } \therefore T_{\mu\nu}^{\mu\nu} = 0$$

$$\therefore (aR^{\mu\nu} + b g^{\mu\nu} R)_{;\mu\nu} = 0$$

Also from Bianchi identity

$$(R^{\mu\nu} - \frac{1}{2} g^{\mu\nu} R)_{;\mu\nu} = 0$$

$$\therefore \underbrace{(aR^{\mu\nu} - \frac{a}{2} g^{\mu\nu} R)_{;\mu\nu}}_{=0} + (\frac{a}{2} g^{\mu\nu} R + b g^{\mu\nu} R)_{;\mu\nu} = 0$$

$$\Rightarrow (\frac{a}{2} + b) (g^{\mu\nu} R)_{;\mu\nu} = 0$$

In general $(g^{\mu\nu} R)_{;\mu\nu} \neq 0$

$$\therefore \frac{a}{2} + b = 0 \Rightarrow b = -\frac{a}{2}$$

$$\therefore a(R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R) = C' T_{\mu\nu} \quad \text{let } C = \frac{C'}{a}$$

$$\therefore R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R = C T_{\mu\nu}$$

Following Chapter 6 on the Notes we see that to recover Newtonian gravity in weak curvature, $C = -8\pi G$

$$\therefore R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R = -8\pi G T_{\mu\nu}$$

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3. a)

$$R_{\lambda\mu\nu\kappa} = \frac{1}{2} \left(\frac{\partial^2 h_{\lambda\nu}}{\partial x^\kappa \partial x^\mu} - \frac{\partial^2 h_{\rho\nu}}{\partial x^\kappa \partial x^\lambda} - \frac{\partial^2 h_{\lambda\kappa}}{\partial x^\mu \partial x^\nu} + \frac{\partial^2 h_{\mu\kappa}}{\partial x^\nu \partial x^\lambda} \right)$$

$$h_{\mu\nu} = A_{\mu\nu} \exp(ik_\rho x^\rho)$$

$$\begin{aligned} \therefore R_{\lambda\mu\nu\kappa} &= \frac{1}{2} \left((ik_\kappa)(ik_\rho) h_{\lambda\nu} - (ik_\kappa)(ik_\lambda) h_{\mu\nu} - (ik_\mu)(ik_\nu) h_{\lambda\kappa} \right. \\ &\quad \left. + (ik_\nu)(ik_\lambda) h_{\mu\kappa} \right) \\ &= \frac{1}{2} \left(-k_\kappa k_\rho h_{\lambda\nu} + k_\kappa k_\lambda h_{\mu\nu} + k_\mu k_\nu h_{\lambda\kappa} - k_\nu k_\lambda h_{\mu\kappa} \right) \end{aligned}$$

linear vacuum field equation

$$R^{(\mu)}_{\nu} = 0 \Rightarrow$$

$$\frac{1}{2} \left(\frac{\partial^2 h}{\partial x^\mu \partial x^\nu} - \frac{\partial^2 h^\lambda_{\lambda}}{\partial x^\mu \partial x^\lambda} - \frac{\partial^2 h^\lambda_{\nu}}{\partial x^\mu \partial x^\lambda} + \square h_{\mu\nu} \right) = 0$$

$$\therefore \cancel{\frac{\partial^2 h}{\partial x^\mu \partial x^\nu}} - k_\mu k_\nu h + k_\nu k_\lambda h^\lambda_{\nu} + k_\nu k_\lambda h^\lambda_{\nu} - k_\rho k^\rho h_{\mu\nu} = 0$$

$$\Rightarrow \cancel{0 = k_\mu k_\lambda h^\lambda_{\nu}} \quad \because k_\lambda h^\lambda_{\nu} = k^\lambda h_{\lambda\nu}, \quad k_\nu k^\lambda = \eta_{\lambda\mu} k^\nu k^\lambda = k^\mu k_\mu$$

$$\therefore \Rightarrow 0 = k_\nu k^\lambda h_{\lambda\nu} - k_\mu k_\nu \frac{h}{2} + k_\mu k^\lambda h_{\lambda\nu} - k_\mu k_\nu \frac{h}{2} - k^2 h_{\mu\nu}$$

$$= k_\nu k^\lambda h_{\lambda\nu} - k_\nu \frac{h}{2} + k_\mu k^\lambda h_{\lambda\nu} - k_\nu \frac{h}{2} - k^2 h_{\mu\nu}$$

$$= k_\nu k^\lambda h_{\lambda\nu} - k_\nu \frac{h}{2}$$

$$= k_\nu k^\lambda h_{\lambda\nu} = k^2 h_{\mu\nu}$$

$$= k_\nu k^\lambda h_{\lambda\nu} - \eta_{\lambda\nu} k^\lambda k_\nu \frac{h}{2} + k_\mu k^\lambda h_{\lambda\nu} - \eta_{\lambda\nu} k_\mu k^\nu \frac{h}{2} - k^2 h_{\mu\nu}$$

$$= k_\nu k^\lambda (h_{\lambda\nu} - \eta_{\lambda\nu} \frac{h}{2}) + k_\mu k^\lambda (h_{\lambda\nu} - \eta_{\lambda\nu} \frac{h}{2}) - k^2 h_{\mu\nu}$$

$$= k_\nu k^\lambda \bar{h}_{\lambda\nu} + k_\mu k^\lambda \bar{h}_{\lambda\nu} - k^2 h_{\mu\nu}$$

$$\Rightarrow \underline{k_x k^\rho \bar{h}_{\rho\nu} + k_\mu k^\rho \bar{h}_{\rho x} - k^2 h_{\mu x} = 0}$$

$\lambda \rightarrow \rho$
 $\nu \rightarrow x$

b) if $k^2 \neq 0$, $h_{\mu x} = \frac{1}{k^2} (k_x k^\rho \bar{h}_{\rho\nu} + k_\mu k^\rho \bar{h}_{\rho x})$

$$\therefore R_{\lambda\mu\nu x} = \frac{1}{2k^2} \left[-k_x k_\mu k_\lambda k^\rho \bar{h}_{\rho\nu} - k_x k_\mu k_\nu k^\rho \bar{h}_{\rho x} \right.$$

$$+ k_x k_\lambda k_\mu k^\rho \bar{h}_{\rho\nu} + k_x k_\lambda k_\nu k^\rho \bar{h}_{\rho\nu}$$

$$+ k_\mu k_\nu k_\lambda k^\rho \bar{h}_{\rho x} + k_\mu k_\nu k_x k^\rho \bar{h}_{\rho x}$$

$$\left. - k_\nu k_\lambda k_\mu k^\rho \bar{h}_{\rho x} - k_\nu k_\lambda k_x k^\rho \bar{h}_{\rho\nu} \right] = 0$$

wave no curvature

so no energy carried.

just wavy coordinates.

$$= k_\nu k^\lambda h_{\lambda\nu} - \eta_{\lambda\nu} k^\lambda k_\nu \frac{\hbar}{2} + k_\mu k^\lambda h_{\lambda\nu} - \eta_{\lambda\nu} k_\mu k^\nu \frac{\hbar}{2} - k^2 h_{\mu\nu}$$

$$= k_\nu k^\lambda (h_{\lambda\nu} - \eta_{\lambda\nu} \frac{\hbar}{2}) + k_\mu k^\lambda (h_{\lambda\nu} - \eta_{\lambda\nu} \frac{\hbar}{2}) - k^2 h_{\mu\nu}$$

$$= k_\nu k^\lambda \bar{h}_{\lambda\nu} + k_\mu k^\lambda \bar{h}_{\lambda\nu} - k^2 h_{\mu\nu}$$

$$\Rightarrow \underline{k_\alpha k^\rho \bar{h}_{\rho\mu} + k_\mu k^\rho \bar{h}_{\rho\alpha} - k^2 h_{\mu\alpha} = 0}$$

$\lambda \rightarrow \rho$
 $\nu \rightarrow \alpha$

5/5

b) if $k^2 \neq 0$, $h_{\mu\alpha} = \frac{1}{k^2} (k_\alpha k^\rho \bar{h}_{\rho\mu} + k_\mu k^\rho \bar{h}_{\rho\alpha})$

$$\therefore R_{\lambda\mu\nu\kappa} = \frac{1}{2k^2} \left[-k_\alpha k_\mu k_\nu k^\rho \bar{h}_{\rho\alpha} - k_\alpha k_\nu k_\mu k^\rho \bar{h}_{\rho\alpha} \right.$$

$$+ k_\alpha k_\lambda k_\mu k^\rho \bar{h}_{\rho\nu} + k_\alpha k_\lambda k_\nu k^\rho \bar{h}_{\rho\mu} \left. \right]$$

$$+ k_\mu k_\nu k_\alpha k^\rho \bar{h}_{\rho\kappa} + k_\mu k_\nu k_\kappa k^\rho \bar{h}_{\rho\alpha}$$

$$- k_\nu k_\alpha k_\mu k^\rho \bar{h}_{\rho\kappa} - k_\nu k_\alpha k_\kappa k^\rho \bar{h}_{\rho\mu} \left. \right] = 0$$

wave no curvature

no so no energy carried.

just wavy coordinates.

3/3

c) if wave propagating at speed of light,
we have $k^2 = 0$

$$\therefore k_\alpha k^\alpha \bar{h}_{\mu\nu} + k_\nu k^\nu \bar{h}_{\mu\alpha} = 0$$

Now set $\kappa = \mu = \sigma$ (we look at the diagonal
terms of the LHS tensor)

$$\Rightarrow 2k_\sigma k^\sigma \bar{h}_{\sigma\sigma} = 0 \quad (\text{No sum over } \sigma)$$

$\therefore k_\sigma$ is generally non-zero

$$\therefore \underline{\underline{k^\sigma \bar{h}_{\sigma\sigma} = 0}} \rightarrow \text{Harmonic gauge}$$

good

4/4

12/12

gives

$$4. \langle L_{GW} \rangle = \frac{32}{5} \frac{G^4}{c^5} \frac{m_1^2 m_2^2 M}{a^5} \left[\frac{1 + \frac{73}{24} e^2 + \frac{37}{96} e^4}{(1-e^2)^{7/2}} \right]$$

$$E_{GW} = \langle L_{GW} \rangle T \quad \leftarrow \text{orbital period}$$

~~As~~ As orbit approaches parabolic

$$L = a(1-e^2), \quad r(1+\cos\phi) = L, \quad e \rightarrow 1$$

Distance of ~~close~~ closest approach occurs at $\phi = 0$

$$\rightarrow 2b = L, \quad \text{ ~~} e \rightarrow 1 \text{ }~~$$

$$\langle L_{GW} \rangle = \frac{32}{5} \frac{G^4}{c^5} \frac{m_1^2 m_2^2 M}{a^{3/2}} \left[\frac{1 + \frac{73}{24} + \frac{37}{96}}{L^{7/2}} \right]$$

$$= \left(\frac{32}{5}\right) \left(\frac{425}{96}\right) \frac{G^4}{c^5} \frac{m_1^2 m_2^2 M}{a^{3/2}} \frac{1}{(2b)^{7/2}}$$

$$= \frac{85}{3} \frac{G^4}{c^5} \frac{1}{2^{7/2}} \frac{m_1^2 m_2^2 M}{a^{3/2} b^{7/2}}$$

$$\therefore T = 2\pi \sqrt{\frac{a^3}{GM}} = 2\pi \frac{a^{3/2}}{G^{1/2} M^{1/2}}$$

$$\therefore E_{GW} = \langle L_{GW} \rangle T = \frac{85}{3} \frac{2}{2^{7/2}} \frac{G^{7/2}}{c^5} \frac{m_1^2 m_2^2 M^{1/2}}{b^{7/2}}$$

$$= \frac{85\pi\sqrt{2}}{24} \frac{G^{7/2} M^{1/2} m_1^2 m_2^2}{c^5 b^{7/2}} \quad \frac{5}{5}$$